

#### Abstract

Spectrally resolved up and down-welling actinic flux was measured from the NCAR C-130 aircraft by the HIAPER Airborne Radiation Package (HARP-AF). Photolysis frequencies were then retrieved for ~40 species. Surface measurements of  $jNO_2$  were made at the BAO Tower in Erie, CO. We examined the effects of aerosol extinction, particularly in boundary layer aerosols that can significantly reduce surface photochemistry rates

In addition, a new technique for detecting UV absorbing aerosols (e.g. BrC, BC) shows promise for real-time flight detection of smoke. Although fires were notably absent during the FRAPPE time period, some smoke plumes were detected by satellite and in the airborne actinic flux measurements. HARP spectral UV measurements provide the opportunity to determine absorption trends in the UV-B where few other measurements exist. Future is work is required to determine if quantitative spectral absorption retrievals are possible.









The C-130 spiraled over the BAO tower allowing comparison between airborne and surface  $jNO_2$ . Weak aerosol in the boundary layer (see UW HSRL above) resulted in mild photolysis reduction below about 3500 masl. The *in situ* surface AOD = 0.120 at 340 nm (derived from AERONET, 0.8 at 550 nm) and input to TUV to produce a profile shape similar to the measurements.

## **Aerosol impacts on photolysis frequencies** in the Denver metropolitan area

1. National Center for Atmospheric Research (NCAR), Boulder, CO

#### **Actinic flux detectors HIAPER Airborne Radiation Package – Actinic Flux** HARP-AF





- Charged-coupled device monochromator
- Spectrally resolved: 280-680 nm •  $2\pi$  steradian quartz hemispherical optics • Fiber optics : doped-OH fused silica for UV transmission
- Secondary scanning system (SAFS)
- deployed to characterize HARP-AF UV

# 

**Optical Collector** 

#### jNO<sub>2</sub> Filter Radiometers **BAO Tower**

- Metcon, GmbH
- $2\pi$  steradian quartz hemispheric optics Downwelling only
- Total from estimated upwelling







**Photolysis Frequencies**  $j[O_3 \rightarrow O_2 + O(^1D)]$  $j[NO_2 \rightarrow NO+O(^3P)]$  $j[H_2O_2 \rightarrow 2OH]$  $j[HNO_2 \rightarrow OH+NO]$  $j[HNO_3 \rightarrow OH+NO_2]$  $j[CH_2O \rightarrow H+HCO]$  $j[CH_2O \rightarrow H_2+CO]$ j [CH<sub>3</sub>CHO  $\rightarrow$  CH<sub>3</sub>+HCO]  $j[C_2H_5CHO \rightarrow C_2H_5+HCO]$ j [CHOCHO  $\rightarrow$  H<sub>2</sub>+2CO] j [CHOCHO  $\rightarrow$  CH<sub>2</sub>O+CO] j [CHOCHO  $\rightarrow$  HCO+HCO] i [CH3COCHO  $\rightarrow$  CH3CO+HCO] j [CH<sub>3</sub>COCH<sub>3</sub>  $\rightarrow$  CH<sub>3</sub>CO+CH<sub>3</sub>] j [CH<sub>3</sub>OOH  $\rightarrow$  CH<sub>3</sub>O+OH]  $j[CH_3ONO_2 \rightarrow CH_3O+NO_2]$ j[CH<sub>3</sub>COCH<sub>2</sub>CH<sub>3</sub>  $\rightarrow$  CH<sub>3</sub>CO+CH2CH<sub>3</sub>] j [CH<sub>3</sub>CH<sub>2</sub>CH<sub>2</sub>CHO  $\rightarrow$  C<sub>3</sub>H<sub>7</sub>+HCO] j [CH<sub>3</sub>CH<sub>2</sub>CH<sub>2</sub>CHO  $\rightarrow$  C<sub>2</sub>H<sub>4</sub>+CH<sub>2</sub>CHOH]  $j[HO_2NO_2 \rightarrow HO_2 + NO_2]$  $j[HO_2NO_2 \rightarrow OH+NO_3]$ j[CH<sub>3</sub>CH<sub>2</sub>ONO<sub>2</sub>  $\rightarrow$  CH<sub>3</sub>CH<sub>2</sub>O+NO<sub>2</sub>]  $j[Br_2 \rightarrow Br+Br]$ j [BrO  $\rightarrow$  Br+O]  $j[Br_2O \rightarrow \text{products}]$ j [BrCl  $\rightarrow$  Br+Cl] j [HOBr  $\rightarrow$  HO+Br] j [BrONO<sub>2</sub>  $\rightarrow$  Br+NO<sub>3</sub>] j [BrONO<sub>2</sub>  $\rightarrow$  BrO+NO<sub>2</sub>]  $j[Cl_2 \rightarrow Cl+Cl]$ j [ClO  $\rightarrow$  Cl+O(<sup>3</sup>P)] j [CIONO<sub>2</sub>  $\rightarrow$  CI+NO<sub>3</sub>] j [CIONO<sub>2</sub>  $\rightarrow$  CIO+NO<sub>2</sub>] j [BrNO  $\rightarrow$  Br+NO] j [BrONO  $\rightarrow$  BrO+NO]  $j[BrONO \rightarrow Br+NO_2]$ j [BrNO<sub>2</sub>  $\rightarrow$  Br+NO<sub>2</sub>] j [CHBr<sub>3</sub>  $\rightarrow$  Products]  $j[CINO_2 \rightarrow CI+NO_2]$ j [CIONO  $\rightarrow$  CI+NO<sub>2</sub>]  $j[N_2O_5 \rightarrow NO_3 + NO_2]$ j [CH<sub>3</sub>CO(OONO<sub>2</sub>)  $\rightarrow$  CH<sub>3</sub>CO(OO)+NO<sub>2</sub>] j [CH<sub>3</sub>CO(OONO<sub>2</sub>)  $\rightarrow$  CH<sub>3</sub>CO(O)+NO<sub>3</sub>] j [CH<sub>2</sub>=C(CH<sub>3</sub>)CHO  $\rightarrow$  Products] j [CH<sub>3</sub>COCH=CH<sub>2</sub>  $\rightarrow$  Products]

logistic team members.

### S. R. Hall and K. Ullmann

Acknowledgements: NCAR is operated by the Universit Corporation for Atmospheric Research under th sponsorship of the National Science Foundation (NSF) Special thanks to all FRAPPE collaborators, C-130 crew and



MODIS reflectance (true color). An aged smoke plume (Utah) overlaps jNO<sub>2</sub> ratio to cloud-free TUV along the C-130 flight track. Dark/light colors indicate reduction/enhancements primarily due to clouds. The light domain and was detected below the aircraft by spectral strongest smoke encounter is shown as a black line in NE Colorado. analysis, but not by in situ biomass burning chemical signatures.



Spectral upwelling actinic flux over adjacent clear and smoke filled Absorption spectra from filters for biomass burning (levoglucosan >50 regions near a fire plume. The flux ratio indicates enhanced UV ng m–3) and non-biomass burning (levoglucosan <50 ng m–3) periods absorption by organic aerosols in the plume. (Hecobian et al., ACP, 2010, doi:10.5194/acp-10-5965-2010).



Ratio of clear/smoke or clear/cloud under varying conditions indicating Smoke detection along the flight track. Qualitative measure of relative UV extinction by comparison of measured actinic flux to cloud-free consistency in the brown carbon enhanced UV absorption in BB and clouds dominated by scattering. TUV fluxes. Below the black line suggests smoke detection.

#### **Future work**

#### **Real-time analysis questions**

- How robust is the analysis (false positives/negatives)?
- What improvements can improve sensitivity and detection?
- Do nadir and zenith and nadir instruments provide vertical information about layers (above, within or below aircraft)?
- What other detectable impacts (scattering, cloud types, albedo, altitude, chemical absorbers, etc)?





#### **Spectral analysis questions**

- How does aerosol size-distribution affect spectral UV extinction (via scattering)?
- Can spectral analysis retrieve UV spectral absorption?
- Can spectral analysis distinguish BrC from BC?
- Can the retrievals be applied to RTMs to determine flux and photolysis profiles?
- Can the results be applied to regional chemistry models to assess photochemical impacts?